

Impact Hazard Assessment for 2011 AG5

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<http://neo.jpl.nasa.gov/risk/>

2011 AG5 Earth Impact Risk Summary

| | |
|---------------------------------|---------|
| Torino Scale (maximum) | 1 |
| Palermo Scale (maximum) | -1.01 |
| Palermo Scale (cumulative) | -1.00 |
| Impact Probability (cumulative) | 2.0e-03 |
| Number of Potential Impacts | 4 |

| | |
|-----------------------|------------|
| V_{impact} | 14.67 km/s |
| V_{infinity} | 9.55 km/s |
| H | 21.8 |
| Diameter | 0.140 km |
| Mass | 4.1e+09 kg |
| Energy | 1.1e+02 MT |

Analysis based on
210 observations spanning 316.77 days
(2010-Nov-08.629742 to 2011-Sep-21.398727)

all above are mean values
weighted by impact probability

Orbit diagram and elements available [here](#).

These results were computed on Mar 31, 2012

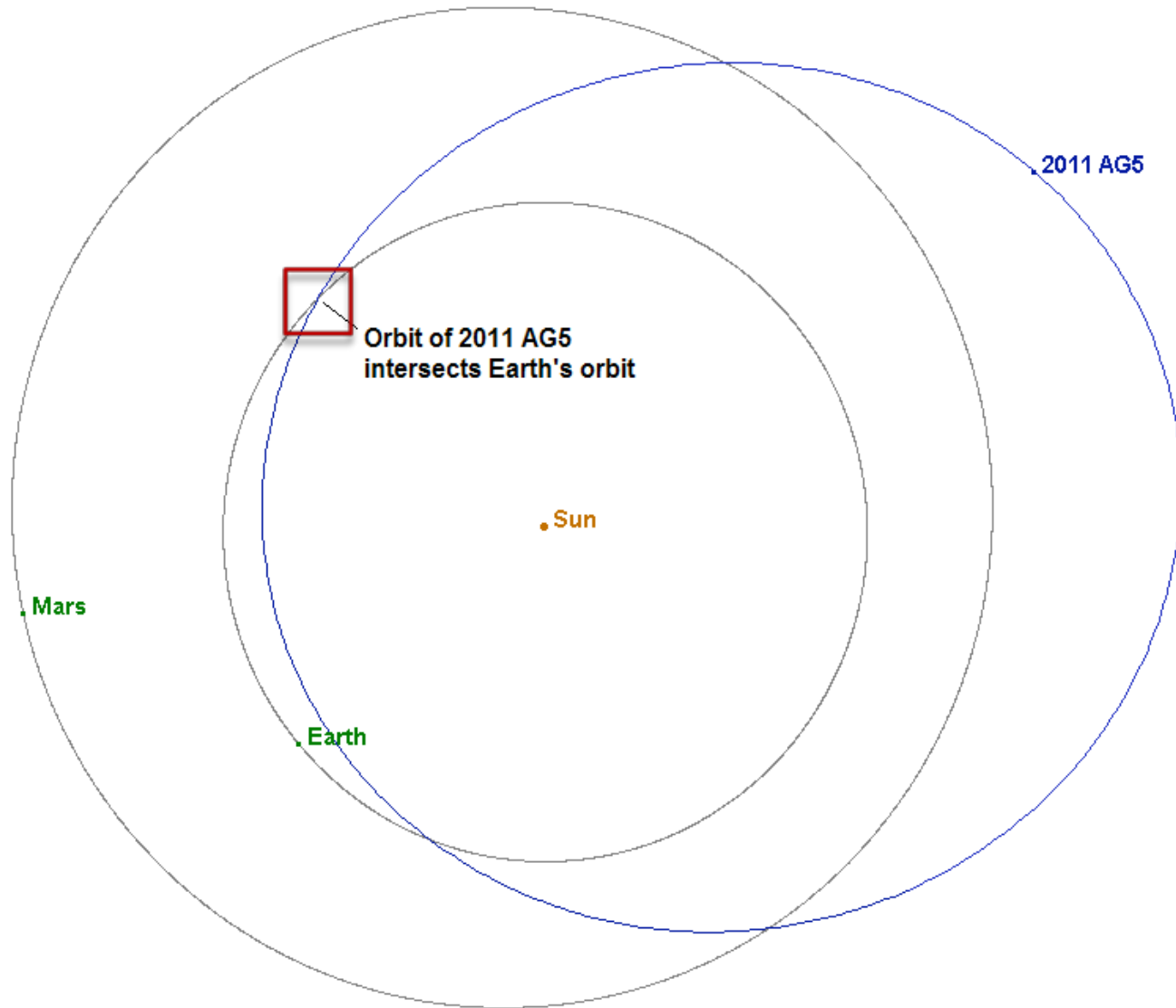
2011 AG5 Earth Impact Table

| Date | Distance | Width | Sigma Impact | Sigma LOV | Stretch LOV | Impact Probability | Impact Energy | Palermo Scale | Torino Scale |
|---------------|------------------------|------------------------|--------------|-----------|------------------------|--------------------|---------------|---------------|--------------|
| YYYY-MM-DD.DD | (r_{Earth}) | (r_{Earth}) | | | (r_{Earth}) | | (MT) | | |
| 2040-02-05.16 | 0.31 | 1.04e-03 | 0.000 | 0.26494 | 3.70e+02 | 2.0e-03 | 1.05e+02 | -1.01 | 1 |
| 2043-02-04.90 | 0.56 | < 1.e-04 | 0.000 | 0.24025 | 1.39e+06 | 4.6e-07 | 1.05e+02 | -4.68 | 0 |
| 2045-02-04.43 | 0.52 | 1.01e-03 | 0.000 | 0.09607 | 5.53e+04 | 1.2e-05 | 1.05e+02 | -3.29 | 0 |
| 2047-02-04.92 | 0.57 | 9.82e-04 | 0.000 | 0.37496 | 1.69e+05 | 3.6e-06 | 1.05e+02 | -3.84 | 0 |

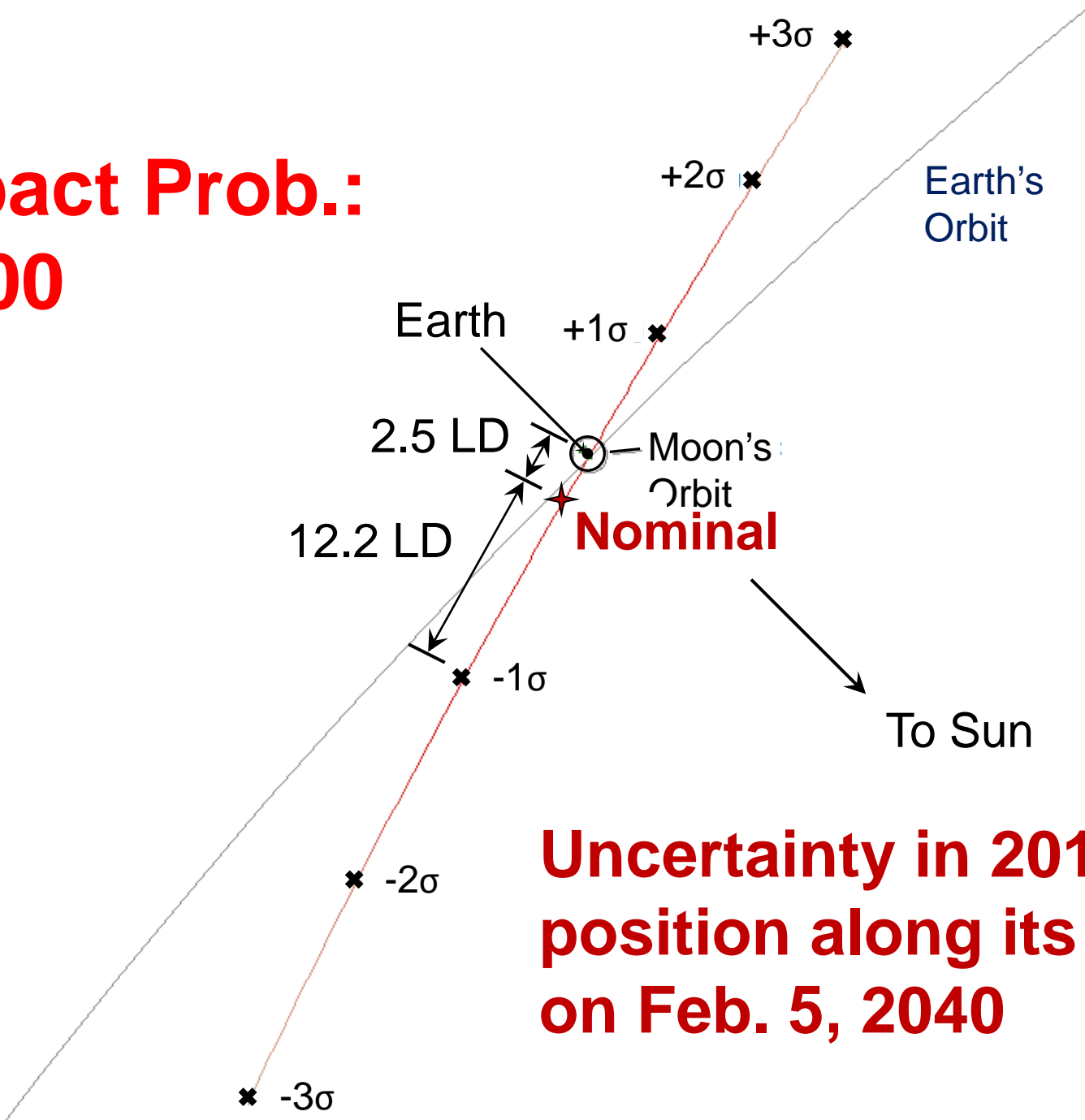
Background - 2011 AG5 in late 2011

- } Discovered Jan. 2011 by Catalina Sky Survey
 - } Pre-discovery obs. by Pan-STARRS dated Nov. 2010
 - } Observations still covered only ~half of 625 day orbit
 - } Asteroid was unobservable for a long period of time
- } Earth impact probability: 1-in-500 for Feb. 5, 2040
 - } Impact requires passage through 365 km keyhole on Feb. 3, 2023
 - } Post-keyhole deflection is ~50x harder than before 2023
- } In 2012, JPL did a full study to answer the key question:
 - } If we wait until AG5 is observable again, and it turns out to be on a collision course, is there enough time to design, build, launch and execute a deflection mission before it passes through the keyhole in 2023, a time span of possibly

Heliocentric Orbit of 2011 AG5

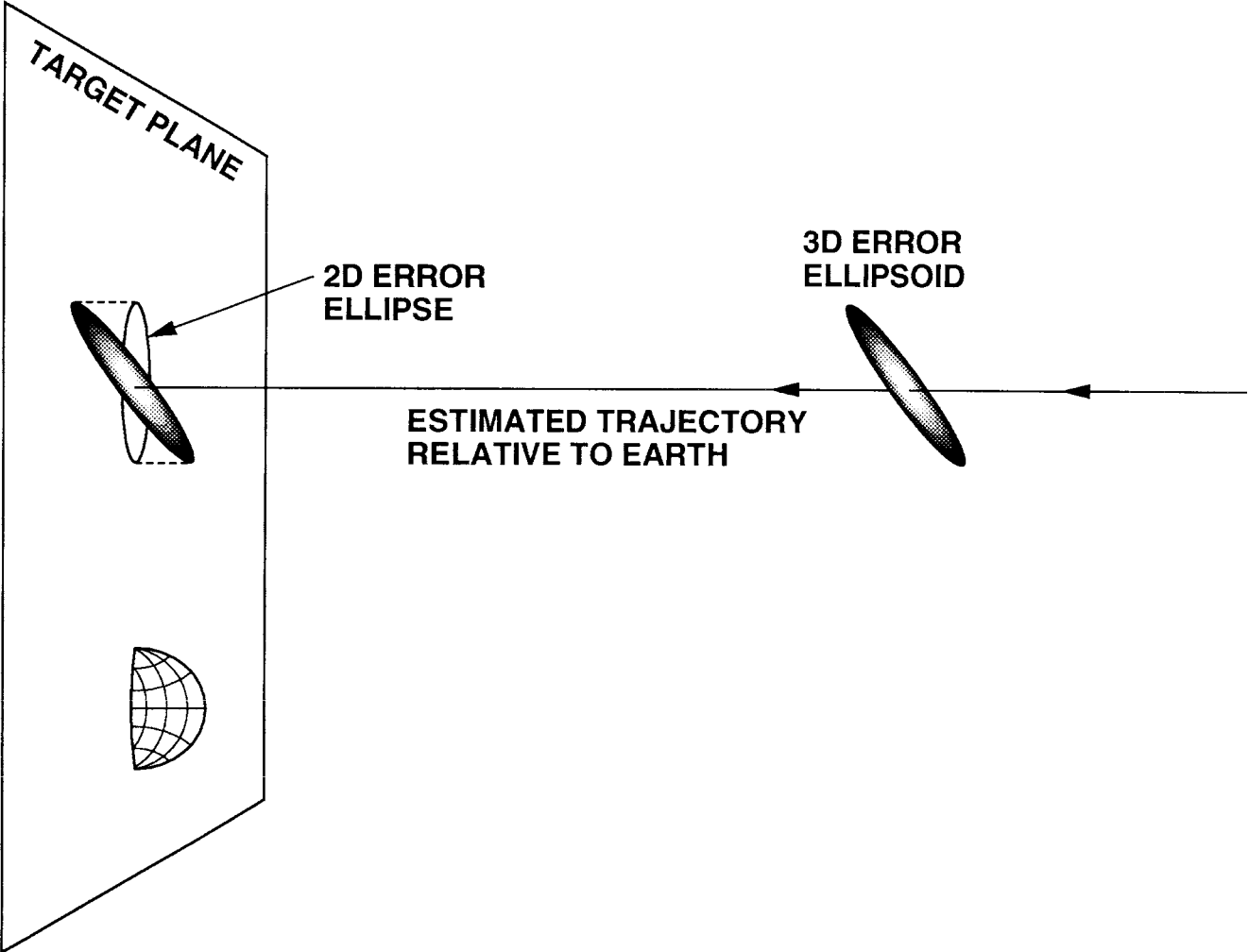


**Impact Prob.:
1/500**

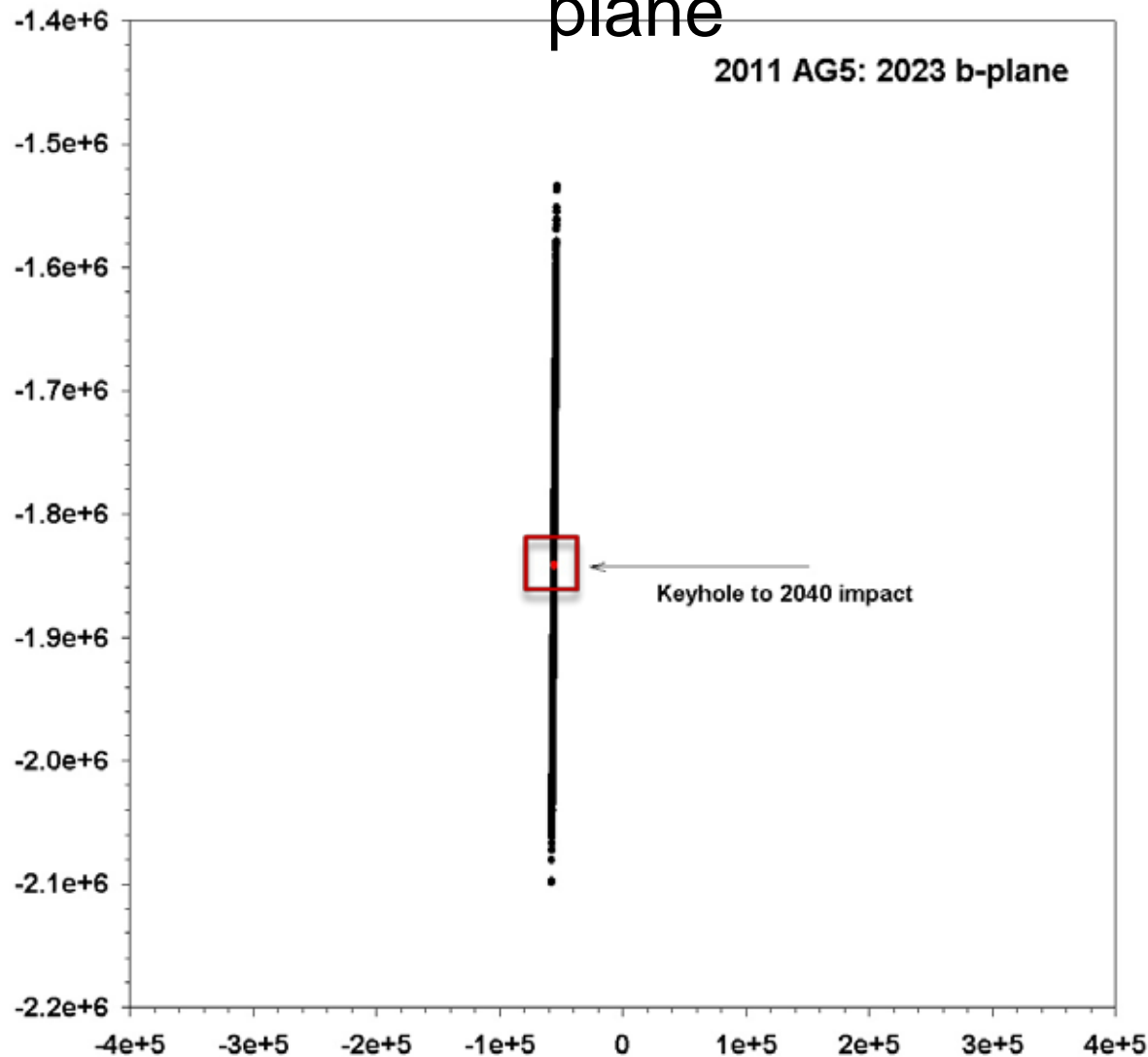


**Uncertainty in 2011 AG5's
position along its orbit
on Feb. 5, 2040**

Uncertainty Region at a Close Approach



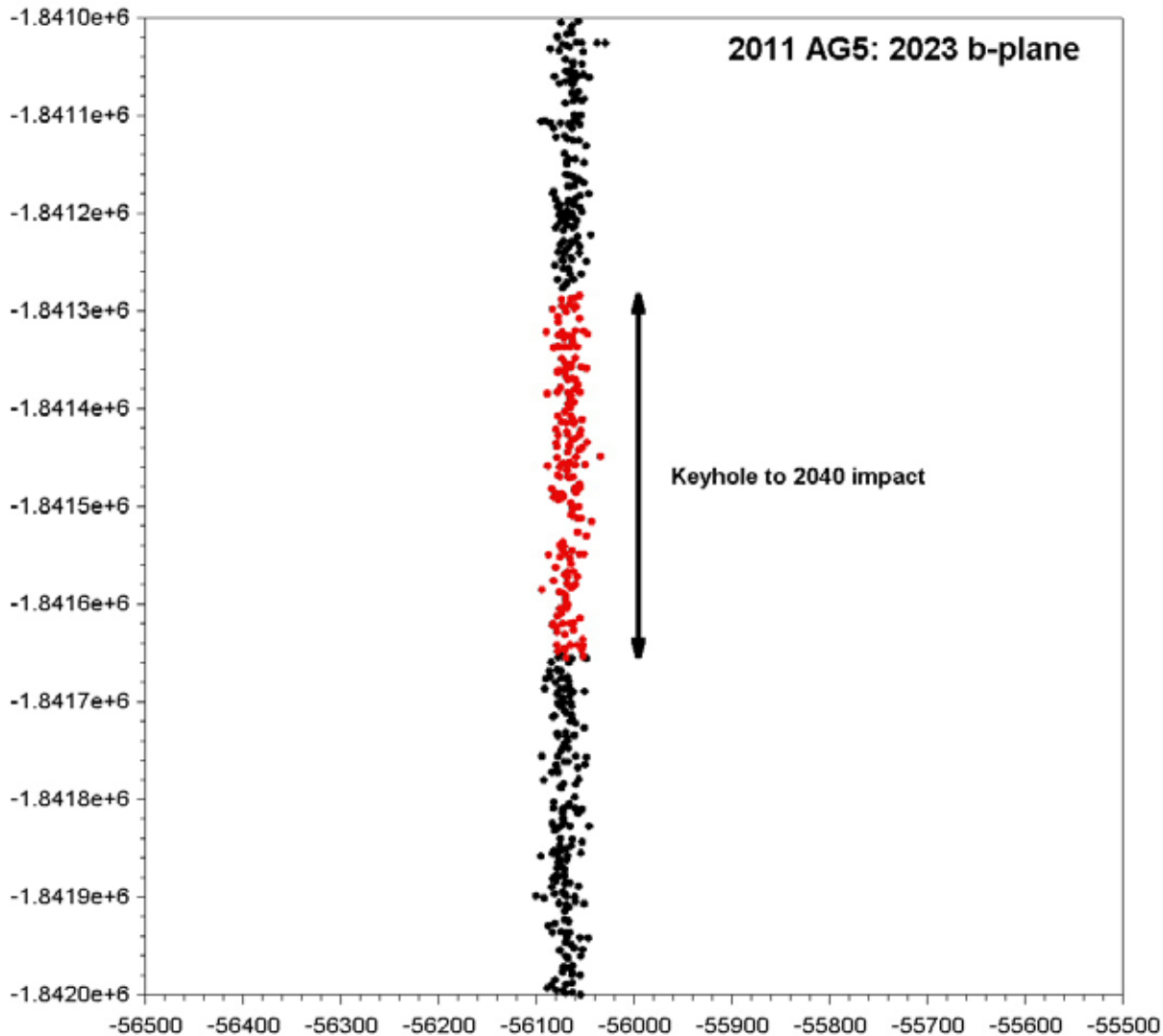
2011 AG5 Uncertainty Region in 2023 b-plane



10,000 Monte Carlo points trace the uncertainty region

Keyhole size:
~365 km

2011 AG5 Keyhole in 2023 b-plane

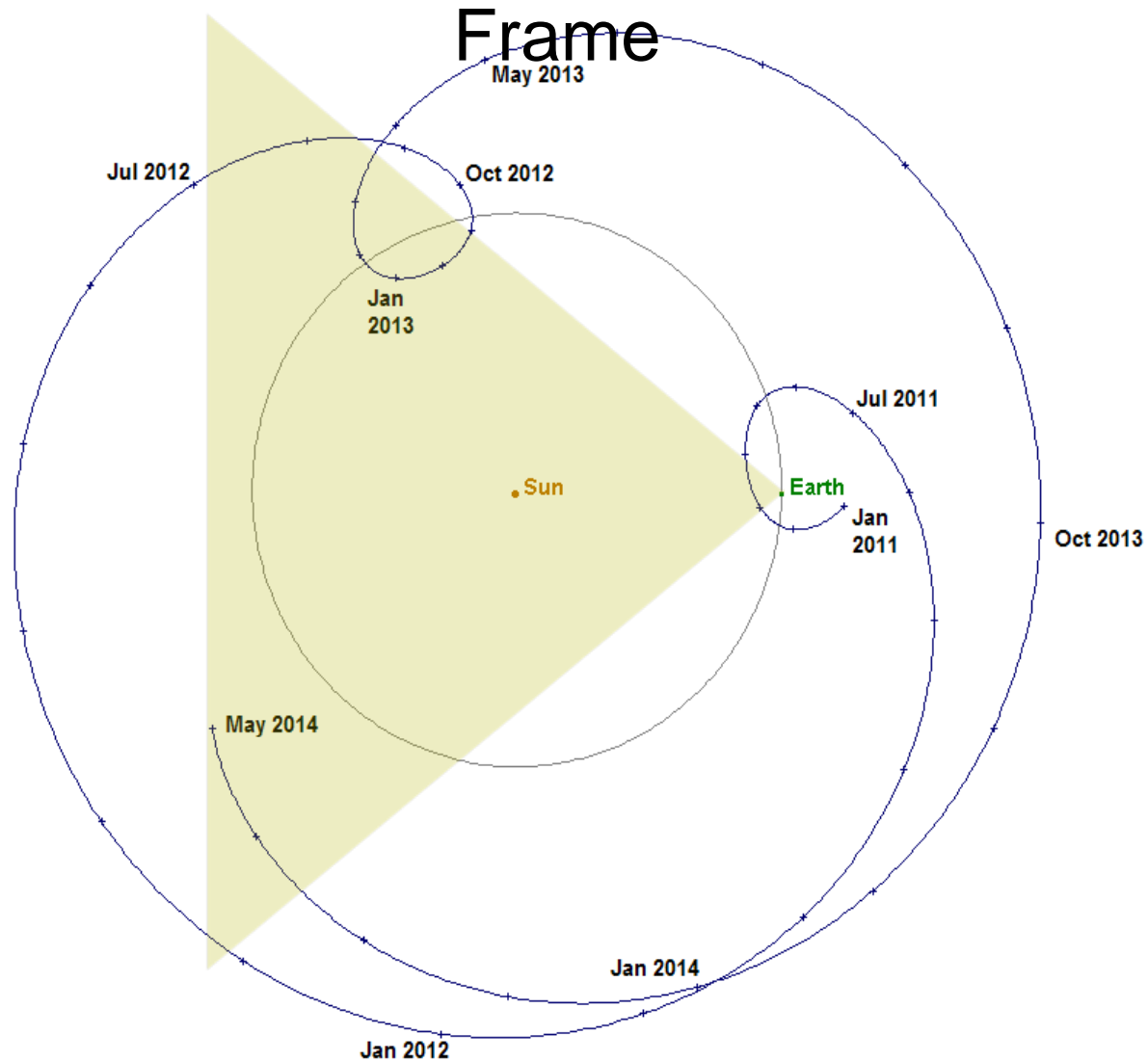


10,000 Monte Carlo points trace the uncertainty region

Keyhole size: ~365 km



Position of 2011 AG5 in a Rotating Reference Frame

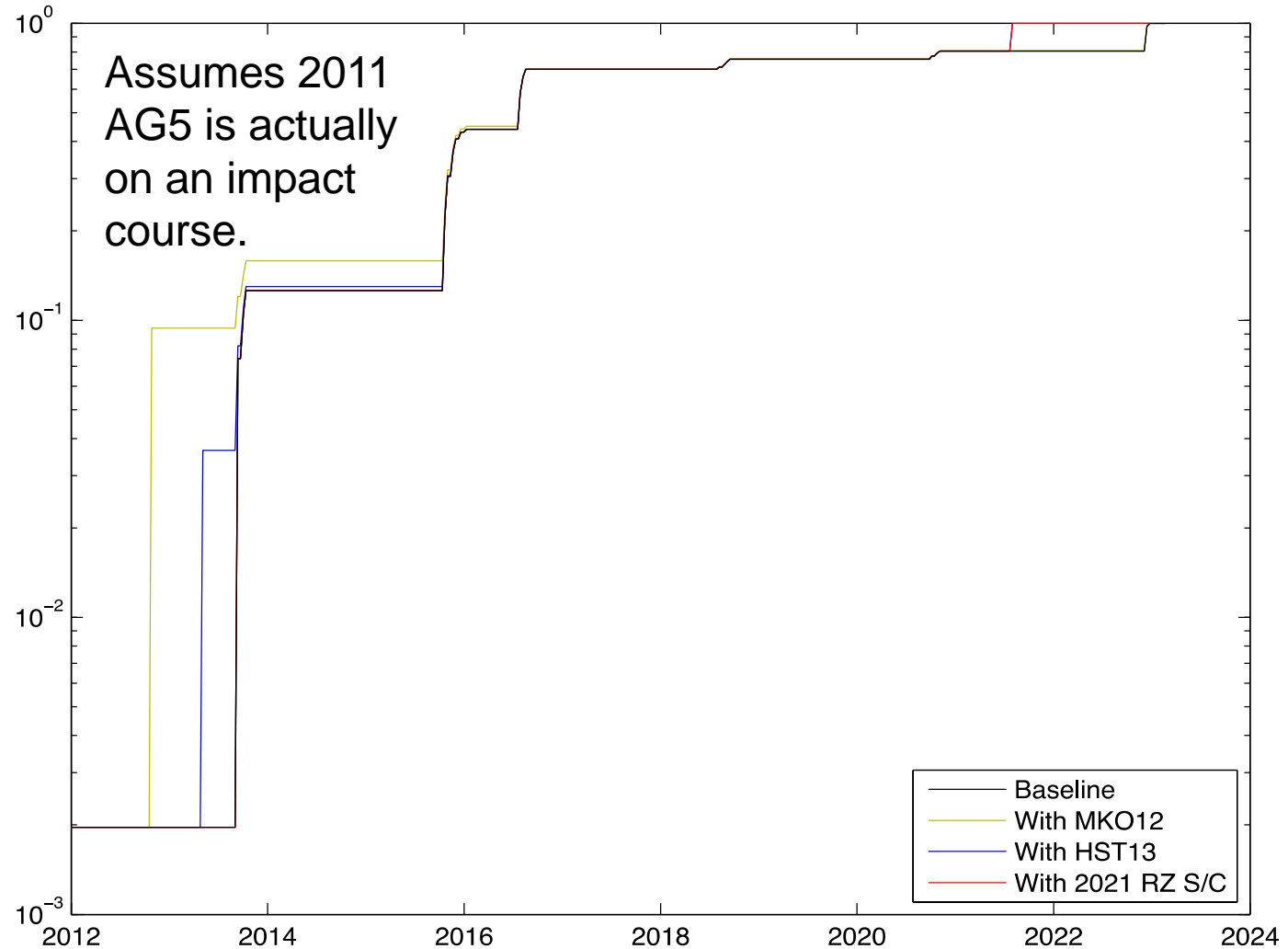


“Future” Observing Opportunities

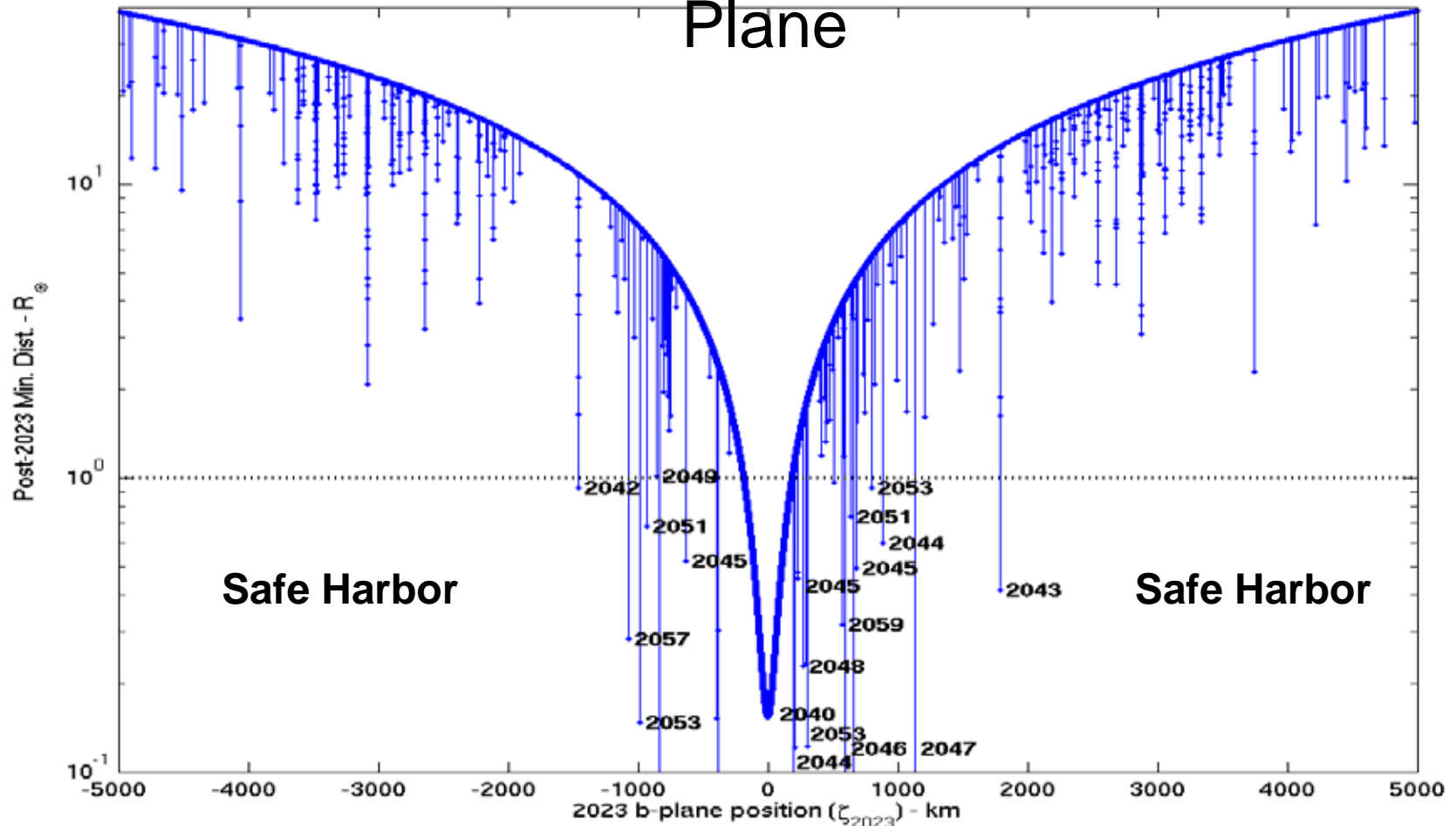
| Date | Brightness (mag) | Solar Elong. (deg) | Remarks |
|-----------|------------------|--------------------|------------------------|
| Oct. 2012 | 24.5 | 42 | Requires Keck. ‘MK012’ |
| Apr. 2013 | 25.5 | 50 | Requires HST. ‘HST13’ |
| Sep. 2013 | 23.6 | 175 | Requires 2-4m aperture |
| Nov. 2015 | 22.9 | 170 | Requires 2-4m aperture |
| June 2016 | 22.9 | 85 | Requires 2-4m aperture |
| Sep. 2018 | 23.1 | 175 | Requires 2-4m aperture |
| Oct. 2020 | 23.5 | 172 | Requires 2-4m aperture |
| Feb. 2023 | 14.3 | 135 | Radar Opportunity |

- } Oct. 2012 obs. require large aperture & favorable conditions
 - } (In fact obtained by Tholen et al. using Gemini 8m and UH 2.2m)
- } April 2013 HST observations
 - } Would require advance characterization of star field
- } “Normal” observations begin in Sept 2013
 - } First observations likely in early August

Maximum Impact Probability vs. Time



Keyhole Map for 2011 AG5 in the 2023 B-Plane



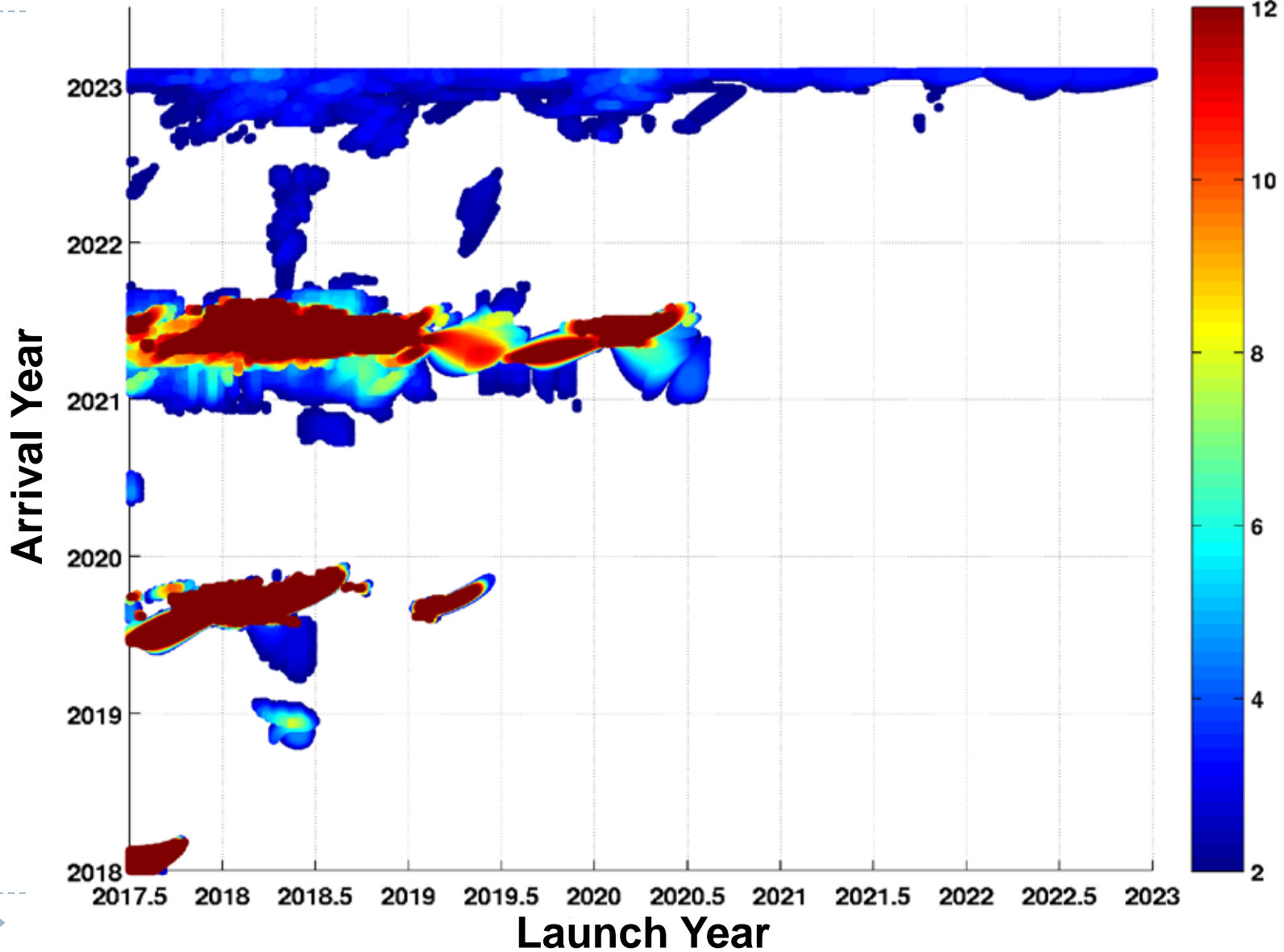
- Secondary keyholes exist but are < 100 m down to a few meters wide.
- Safe harbor zones: -8,000 km to -1,500 km on left & +2,000 km to +12,000 km on right
- ▶ Left safe harbor is preferred because it corresponds to front side impact by S/C

Deflection Campaign (see poster by Damon Landau)

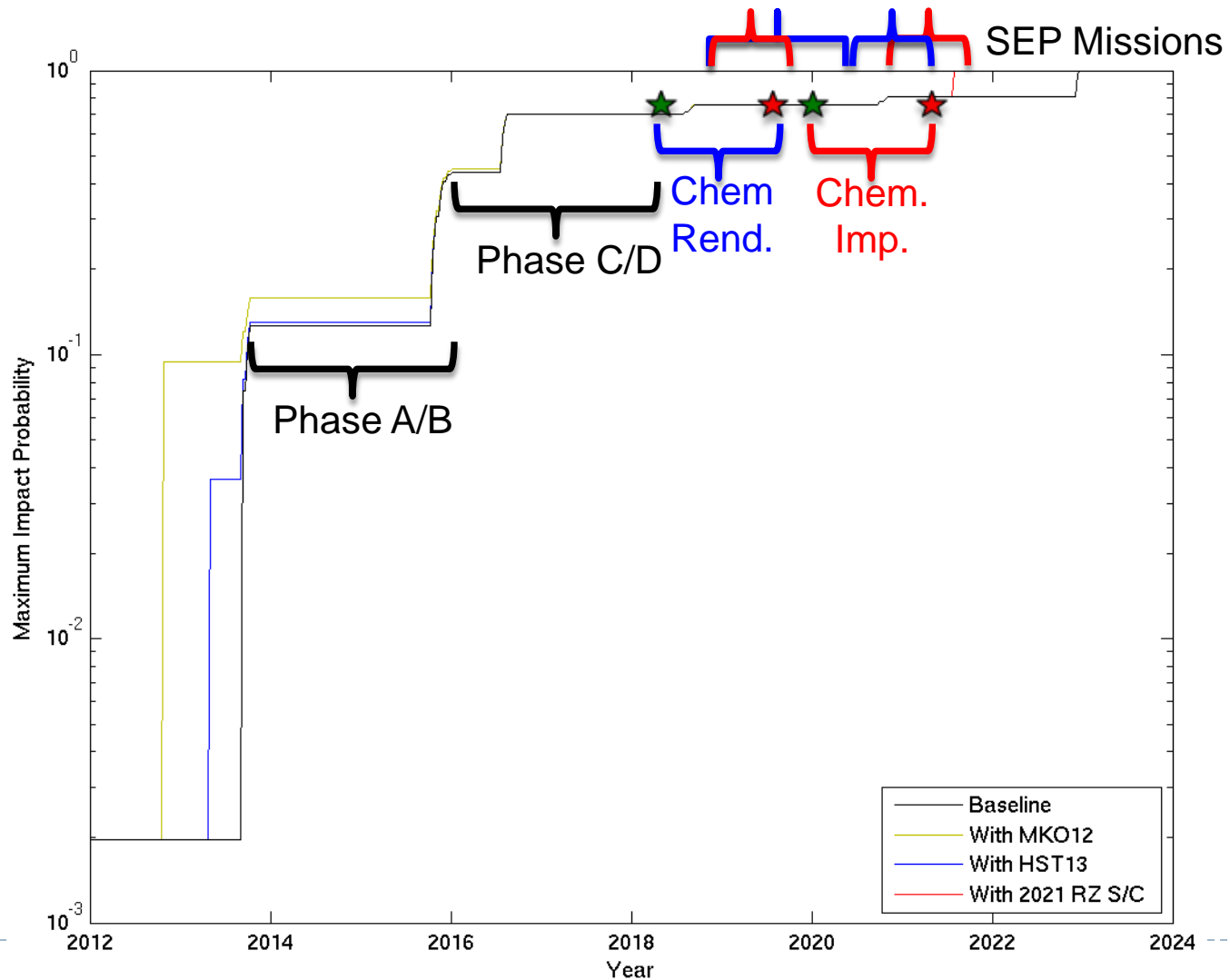
- } Kinetic impactor deflection with observer spacecraft
 - } Examine both chemical & solar electric propulsion (SEP) missions
 - } Require precursor rendezvous spacecraft arriving >2 months before impactor to aid targeting and confirm successful deflection
- } Tune spacecraft mass to obtain the desired deflection
 - } $\Delta V = \beta/M \times V_\infty \times m$
 - } B is the momentum enhancement due to impact ejecta (likely range: 1 to 4)
 - } M is the mass of the asteroid
 - } Take safety factor of 10 on β/M , and so strive for >10 R_E deflection
 - } But if β/M is much higher than expected could lead to a deflection approaching 100 R_E
- } Without early reconnaissance it may be impossible to
 - ▶ ensure that deflection moves asteroid to a “safe harbor” (8-

Pre-Keyhole Deflection Options Atlas V (401)

Impact Δb , Earth radii



Mission Timelines

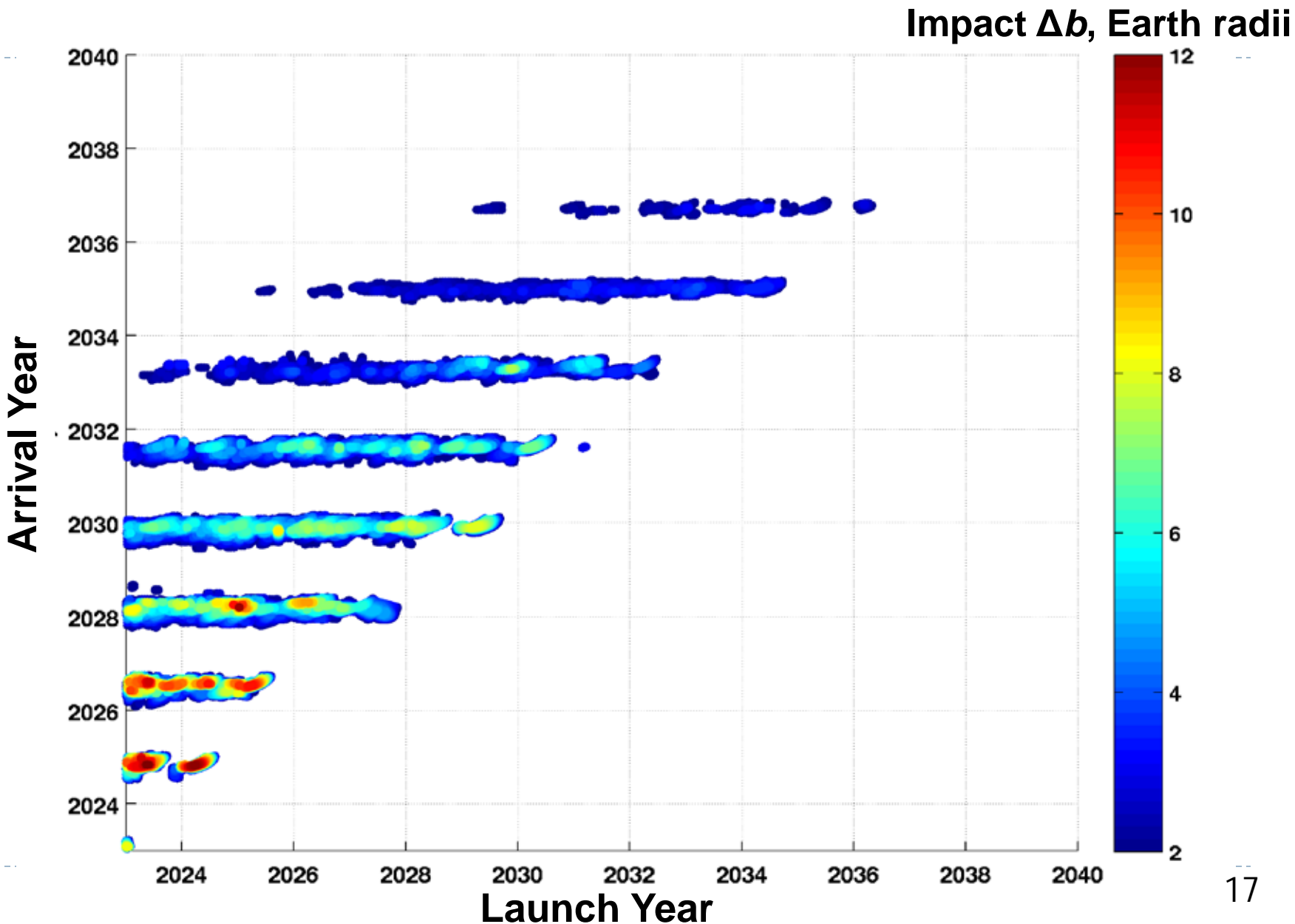


Post Keyhole Mission Designs

- } Post keyhole missions are ~50 more challenging but there are viable rendezvous/deflection options after 2023 that could be carried out with existing launch vehicles
 - } Backup in case pre-keyhole missions unsuccessful
 - } Both chemical and SEP propulsion options are available



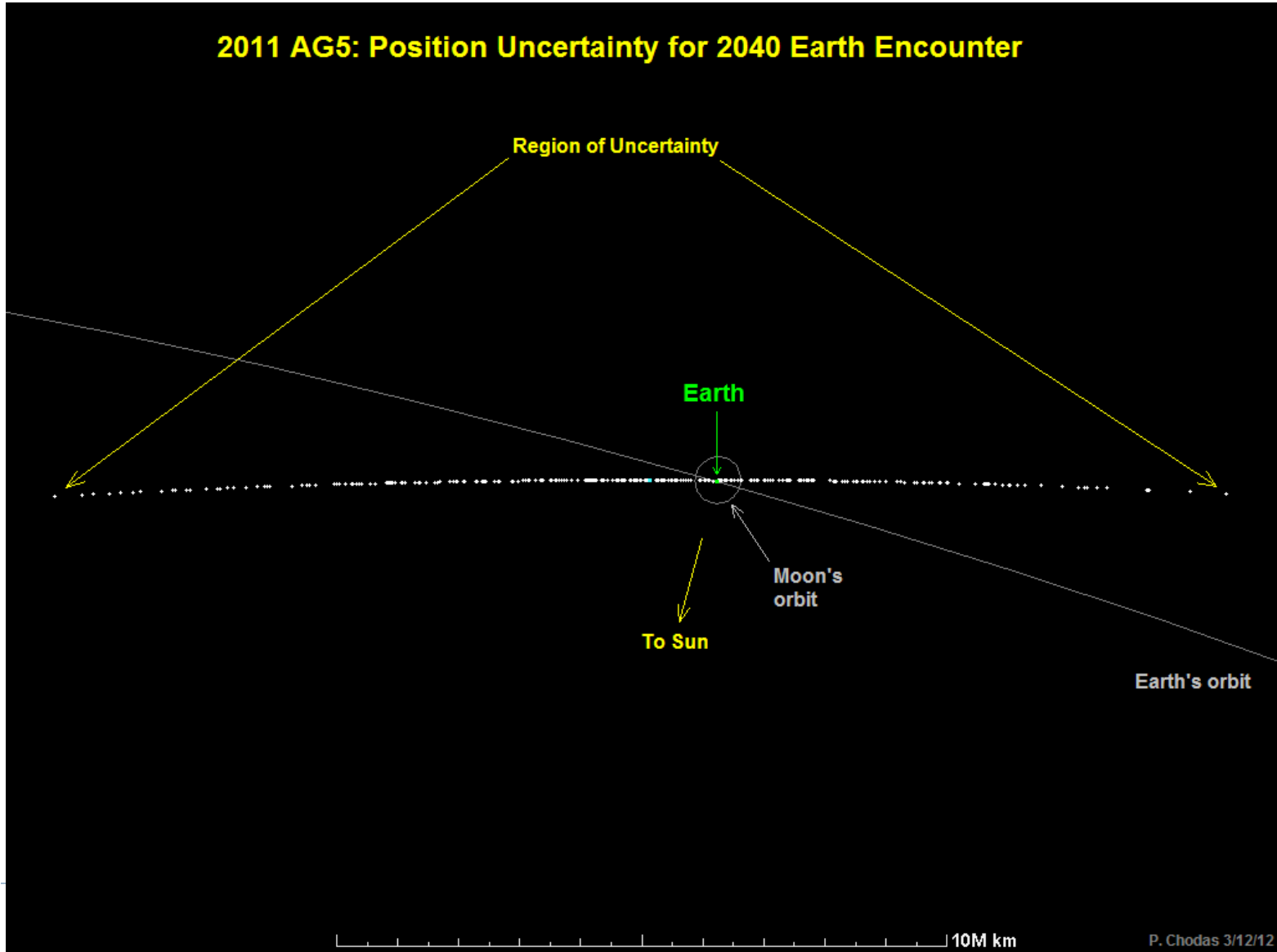
Post-Keyhole Deflection Options (Delta IV Heavy)



Key Conclusions from the 2012 Report:

- } If 2011 AG5 really is on an collision trajectory, the next observations will cause the impact probability to jump to ~10% or more
- } In the unlikely case where the 2012/2013 observations do not eliminate the potential hazard, there is time to plan and carry out a pre-keyhole rendezvous and deflection mission from that point.
- } There exist numerous viable rendezvous/deflection mission options both **before** and **after** keyhole in 2023
- } The full report is available online:
 - } <http://neo.jpl.nasa.gov>

Postscript: Uncertainty Region Before 2012 Obs.



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