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# **APOPHIS 2029** Preliminary results of a CNES feasibility study

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#### **APOPHIS**

- Discovered in June 2004
- Close approach to the Earth on April 13, 2029
- Tumbling spin state with a period ~ 30 hours

Rationale of the CNES study

- priority to the mitigation preparation objectives scientific objectives not excluded but secondary

General features (shape, mass, rotation, gravity) Wide Angle and Narrow Angle cameras in VIS and near IR

Apophis during its 2029 pass?

- first mission to Apophis

Main mission objective

=> Internal structure characterization

Strawman Payload

- Nac and Wac in VIS and Near IR

- Spectroscopie in VIS and IR large spectrum

mission, if needed:

Radio science

Sub-surface analysis

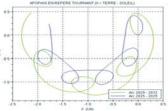
- High Frequency Radar

Surface analysis

Options:

Assumptions:

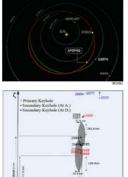
- Diameter ~ 330 meters (from Herschell IR observations)
- Next observation window from the Earth in 2020-2021



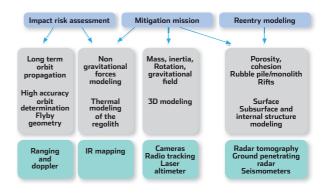
=> Observation of its possible changes induced by the close approach

Is there a significative advantage to plan for a space mission where a probe would accompany

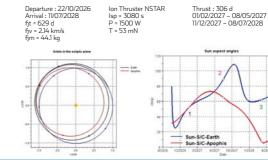
Provide informations about Apophis that would be requested for designing a future mitigation



## Instrumentation / Objectives



### Mission analysis example

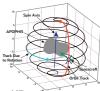


#### **Operational scheme**

- Rendez-vous to a waiting position 200 km in the sun direction (mid July 2028)
- Approach to 20 km controled from the ground using NAC (1m@10km)
- VIS and IR science between 20 km and 3 km
- Radio science between 300 m and 3 km
- About 6 months for the programme before April 13
- Same observation programme after the flyby

#### Main outcomes

- Compatibility with a SOYOUZ launch from French Guyana Space Center
- Electric propulsion required
- Affordable fuel budget (~50 kg)
- Mitigation objectives can be achieved through VIS and IR cameras and 2 radars
- The mitigation objectives can be satisfied without landing any device
- The 2029 timeframe is a unique opportunity to study Apophis, its internal structure, the possible changes due to the gravitational effects from the Earth and determine the exact geometry of its flyby



Thermal map of Apophis from Herschel obser at 70, 100 and 160 microns. @Esa, Herschel, Mach-11, T. Müller, MPE

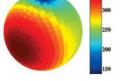


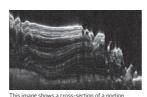
Sismometers
Artificial craterisation

- Long term tracking:
- Radio and/or laser reflector to be assessed

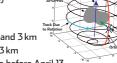
#### Radars

- Low frequency radar for tomography
- Tomography in Reflection • RF wave propagation (~20 MHz)
- Orbiter only (Marsis-like)
- Deep interior profiles : Interfaces, layers and voids
- Heritage Sharad, Marsis
- High frequency radar for regolith characterisation - 300 MHz-1GHz
- 10 cm resolution @10 m depth
- Can be used in an altimeter mode and contribute
- to the gravitational field determination
- Used in the range 500 m -2500 m
- Heritage Wisdom/EXOMARS





This image shows a cross-section of a portion of the north polar ice cap of Mars, derived from data acquired by the Mars Reconnaissance Orbiter's Shallow Radar (SHARAD).



- Spacecraft ahead of Apophis for the flyby
- High accuracy geometry determination of the F/B
  - from ranging/doppler between G/S and S/C (eg from KRU)
- to identify possible changes